

# Overview of the collaborative « Magnetism » program and HEDP magnetic reconnection study

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# Outline

- « Magnetism » program
- Current status of research
- Experimental measurement of B-field produced by high-power laser
- Simulations of reconnexion in non-HEDP and HEDP configurations
- Next stages

# « Magnetism » program

- The program is aimed at improving the current understanding in B-field
  - generation, growth (in relation with heat flow)
  - reconnexion with applications in HEDP (ICF) and space physics
- Collaboration between HELP labs (INRS, LULI) and space physics labs (LPP, LESIA)
- Includes experimental (using lasers) and theoretical effort
- Open to collaborations!

# Past and ongoing effort

- Clarify the interplay between magnetic fields and heat flow in order to have **predictive capability regarding B field growth and evolution in HEDP**
  - B fields affect heat carrying electrons (gyration & heat flow bending through Righi-Leduc effect)
  - Heat flow & non-locality affect B fields through the Nernst effect
- Previous work (IC, MIT, LLE) have underlined the complexity of the topic & the ability of VFP simulations to model correctly such dynamics
- Our work: improve MHD modeling & B fields/heat flow measurements in various HEDP conditions
- Explore the applicability of HEDP experiments to space & solar physics issues, namely reconnexion in various configurations

# A first experiment conducted at RAL clearly showed the toroidal nature of the probed B-fields

INRS  
Université d'avant-garde

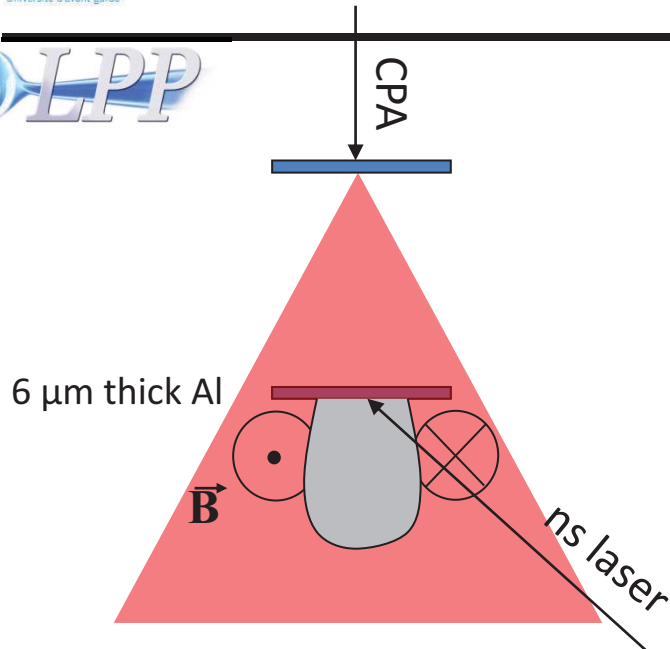
LPP



Observatoire  
de Paris

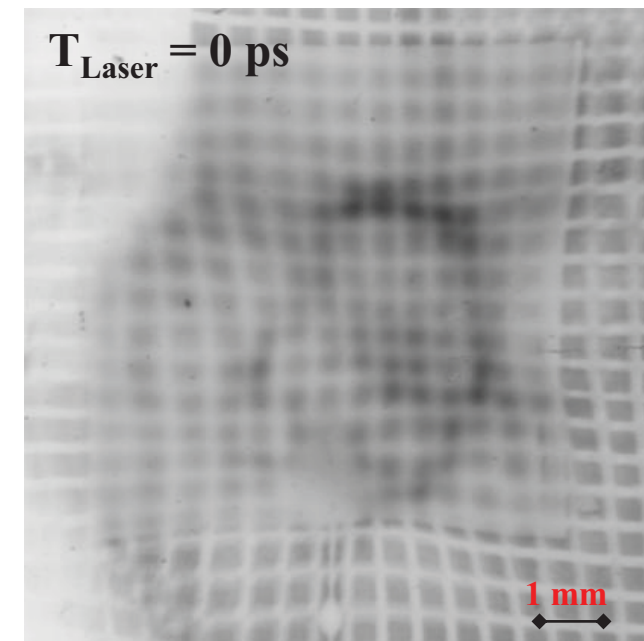
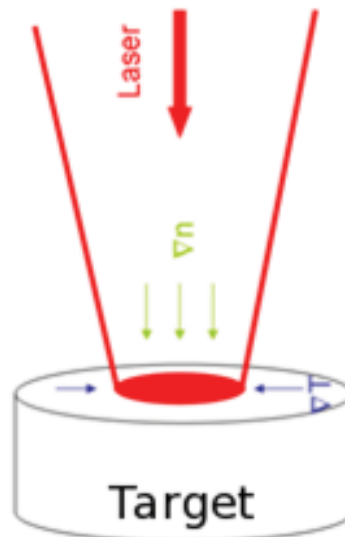
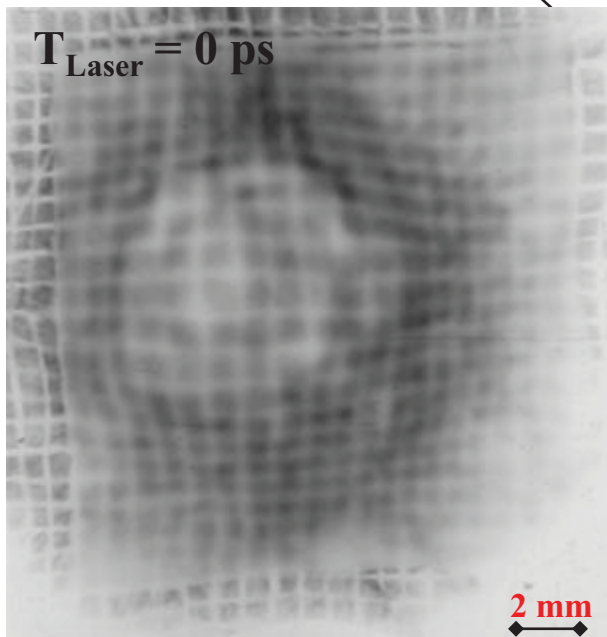
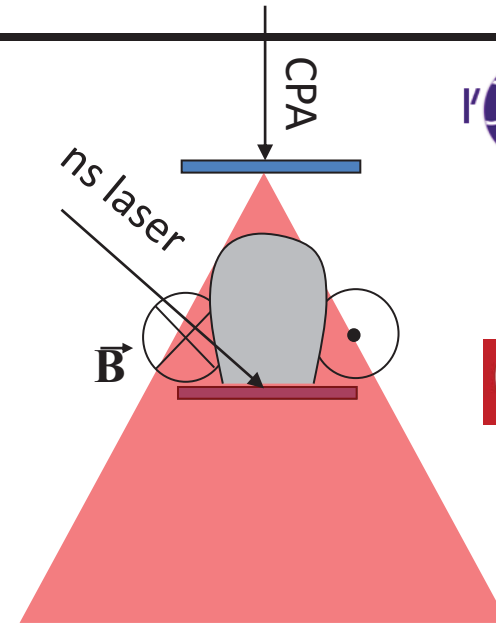


LNCMI

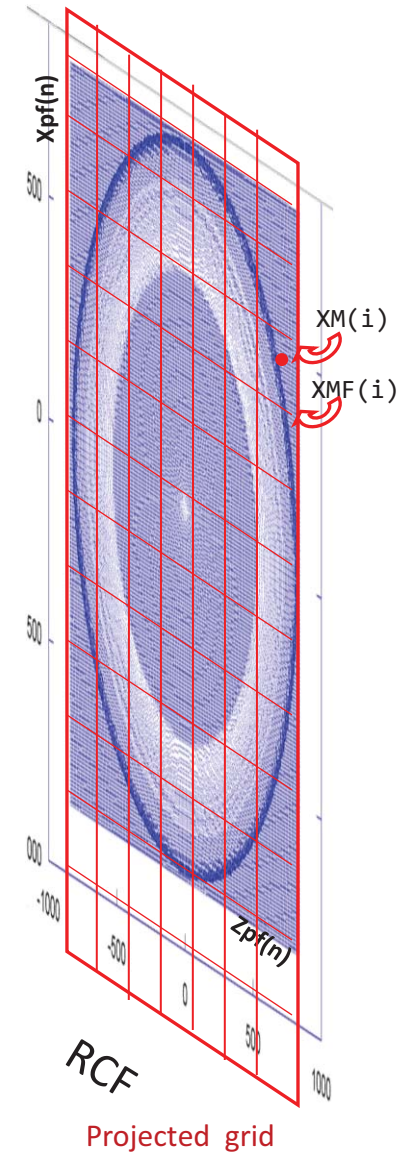
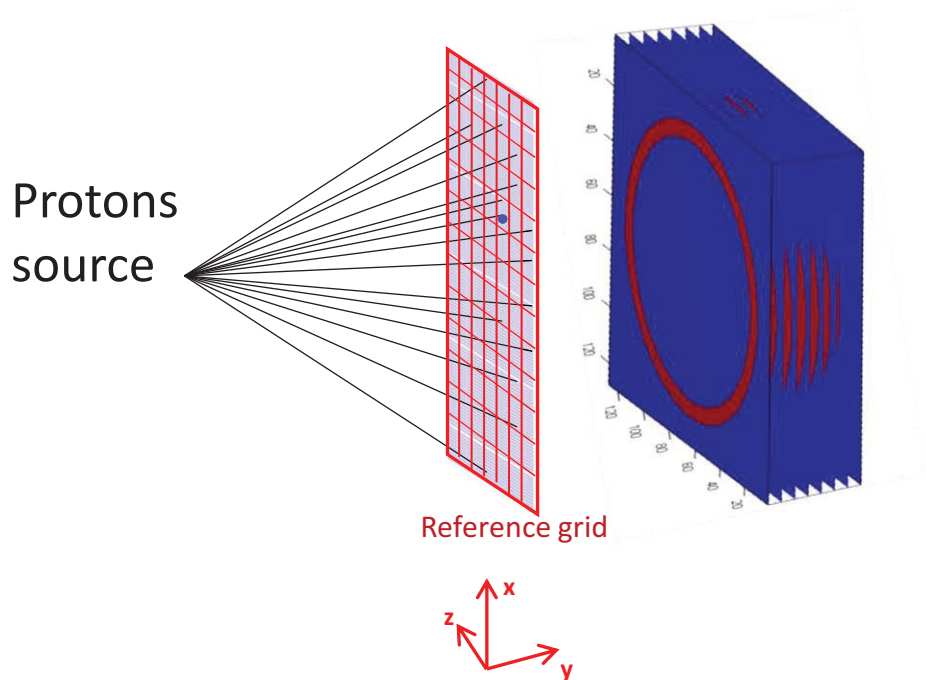


RAL Vulcan  
C. Cecchetti et al., PoP 2009

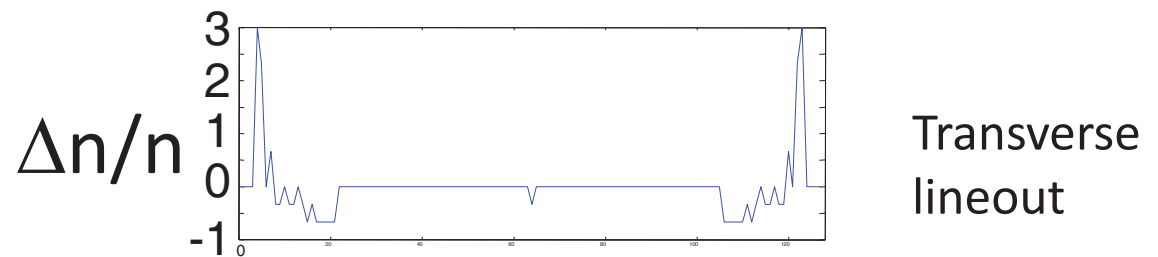
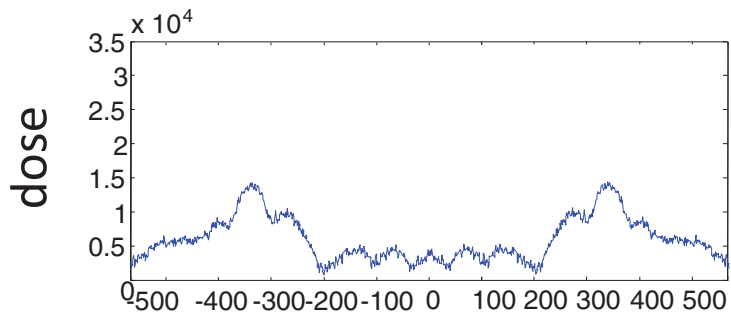
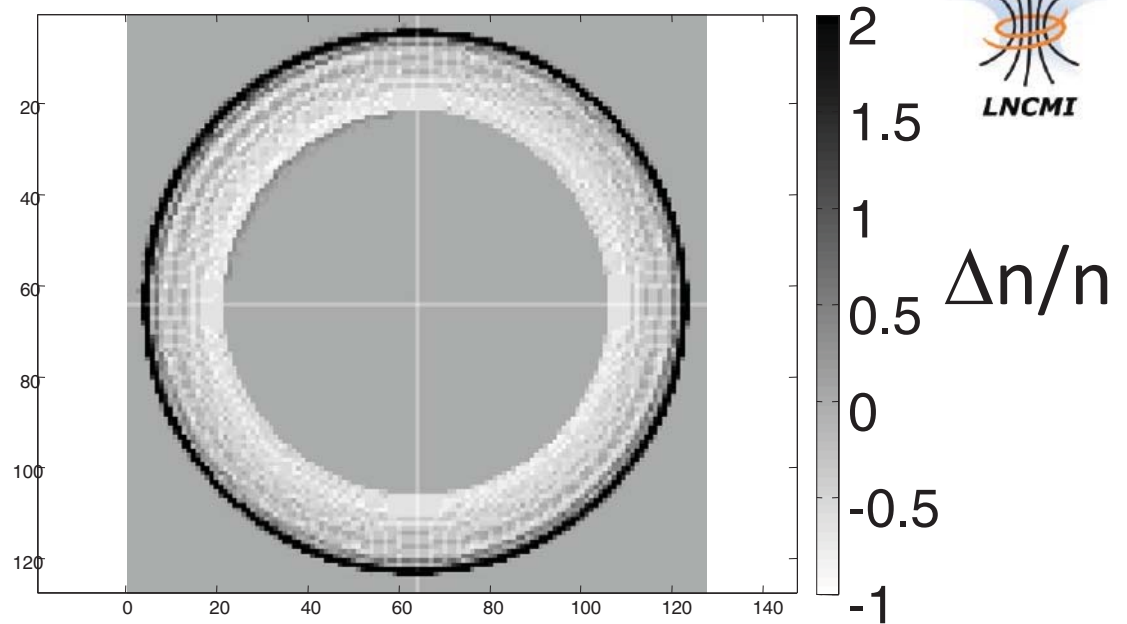
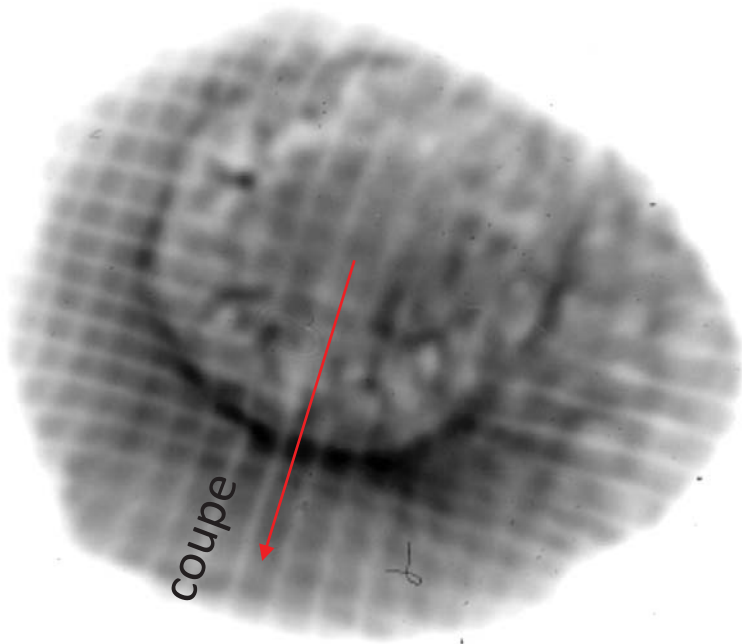
50 J, 1 ns at 1 μm  
focal spot of 50 μm  
 $I = 3-6 \times 10^{14} \text{ W/cm}^2$



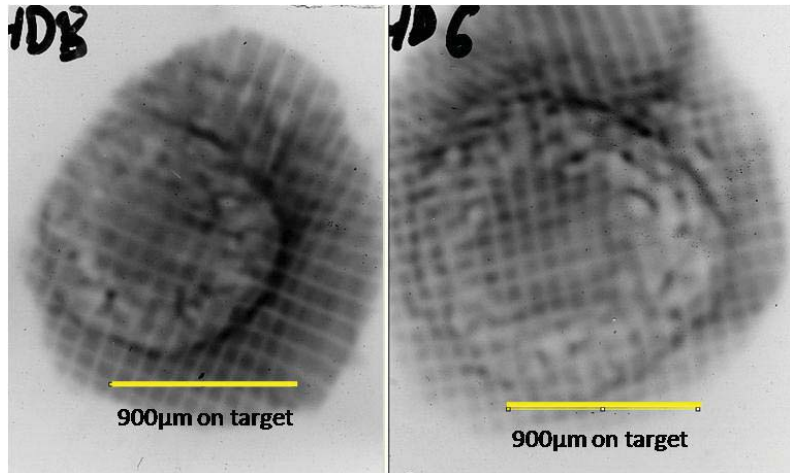
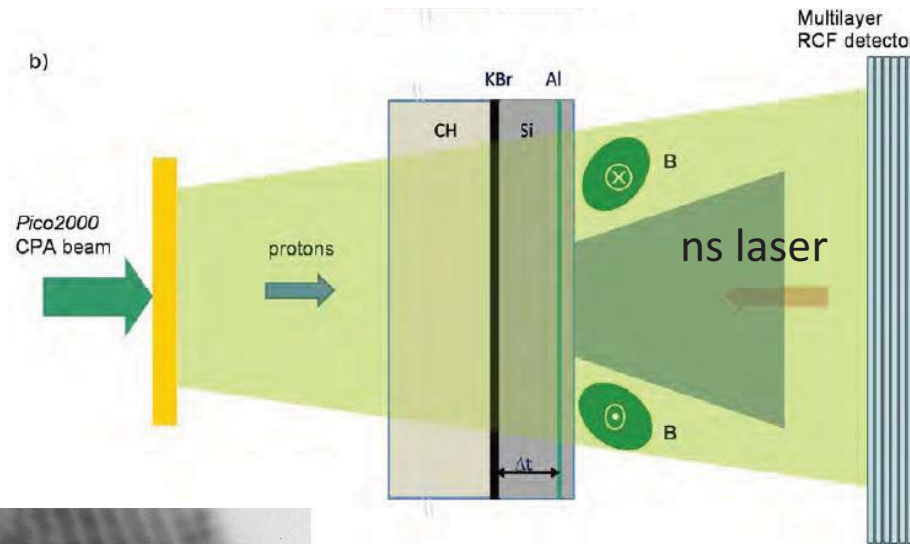
# To allow comparison with hydro-rad simulations, we compute 3D proton trajectories in B-fields



Both the dose modulation and the location of the deflexion are used to compare with the experiments



# A second experiment was performed at LULI2000 using higher energy (200-400 J) and large focal spot

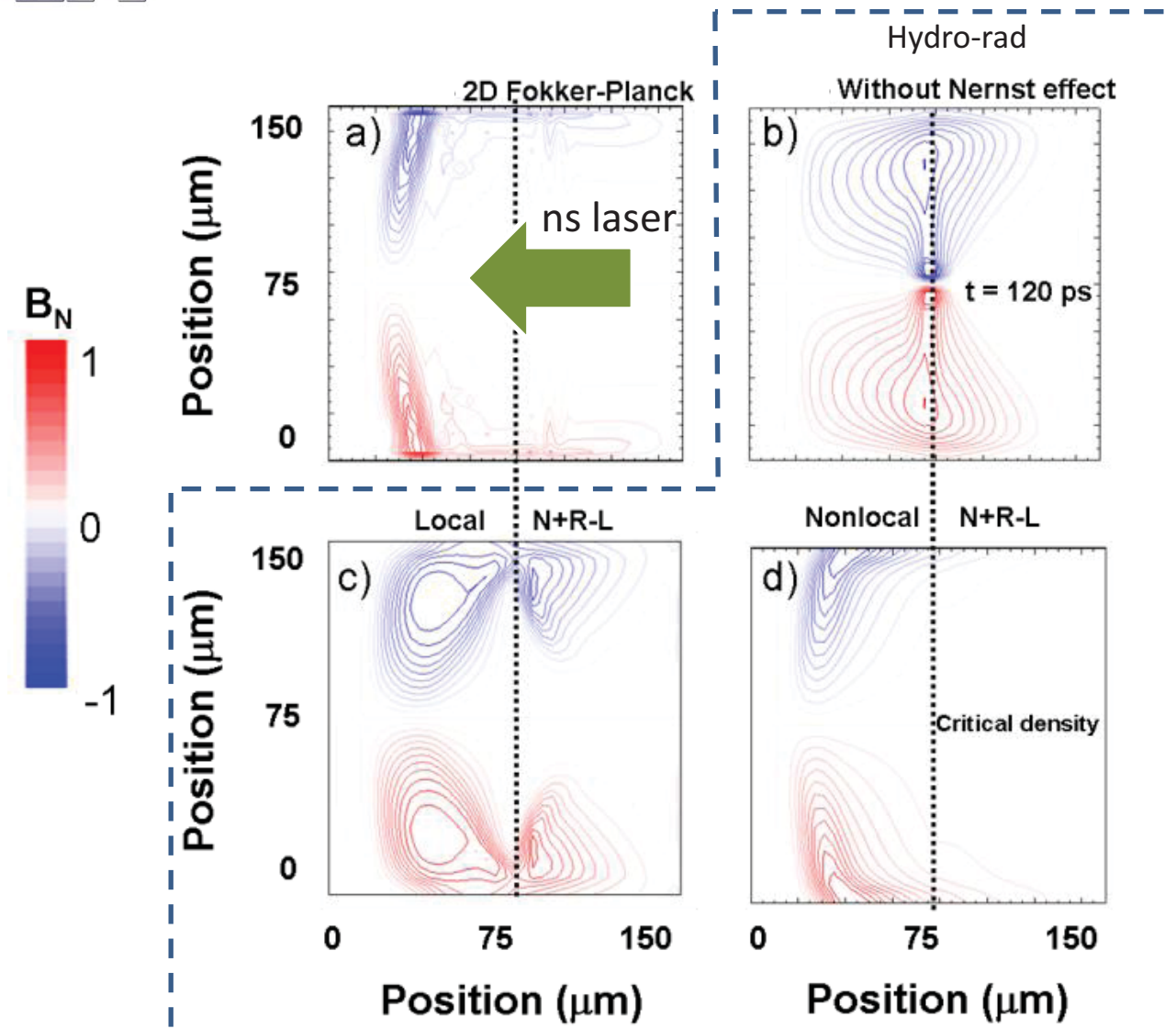


100µm Gaussian      200µm flat top  
probed 0.6 ns after the onset of the  
 $\sim 2 \cdot 10^{14} \text{ W/cm}^2 / 2 \text{ ns}$  beam

- Large deflections were observed
- The longitudinal heat flow measured through x-ray tracers was consistent with the simulations



# Significant efforts were made on the simulation side

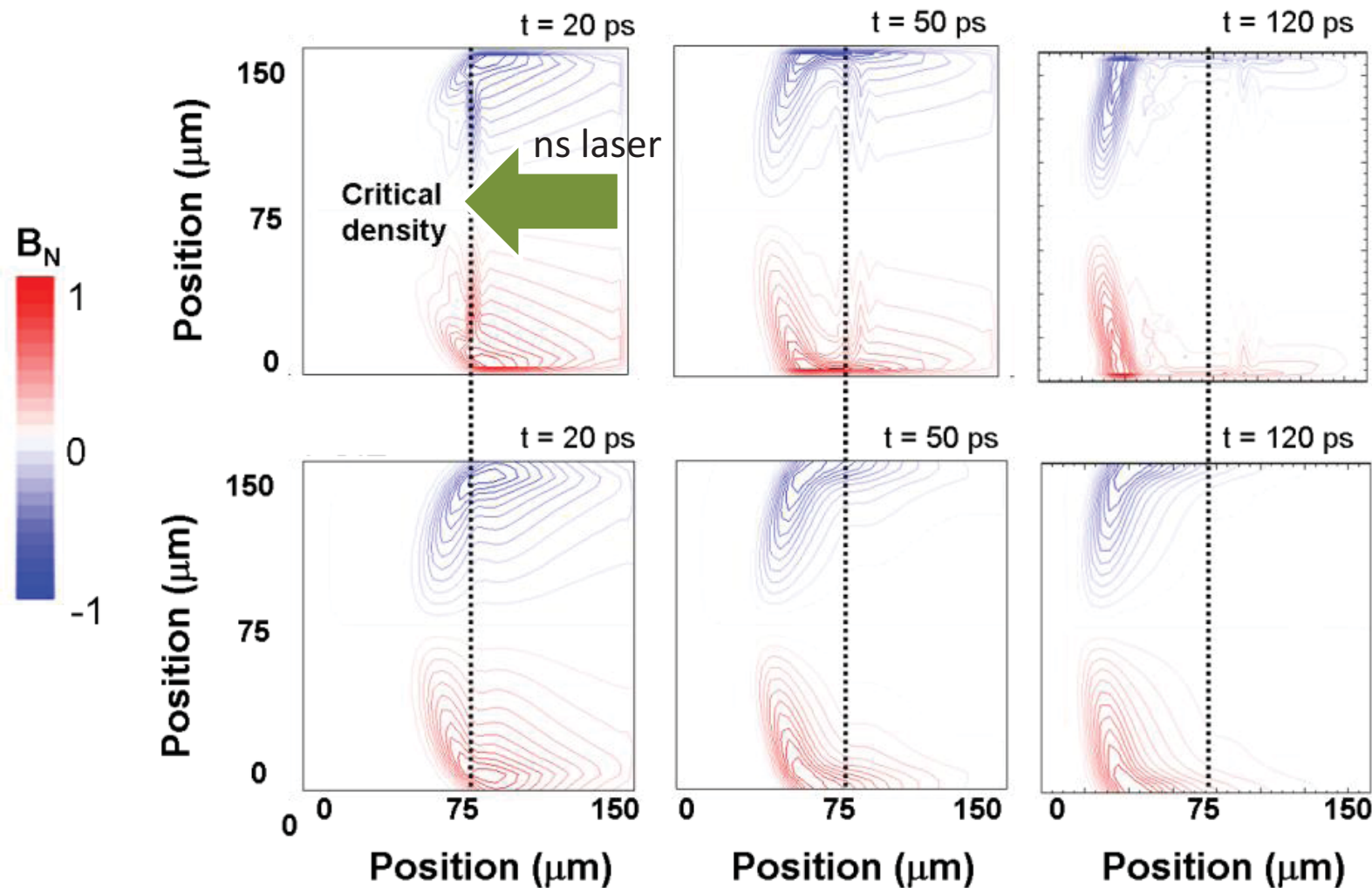


Comparison between:

- kinetic code
- hydro-rad code with detailed MHD

(taking into account pressure driven source term, resistive diffusion of magnetic field and non local electron transport with magnetic inhibition)

# Significant efforts were made on the simulation side



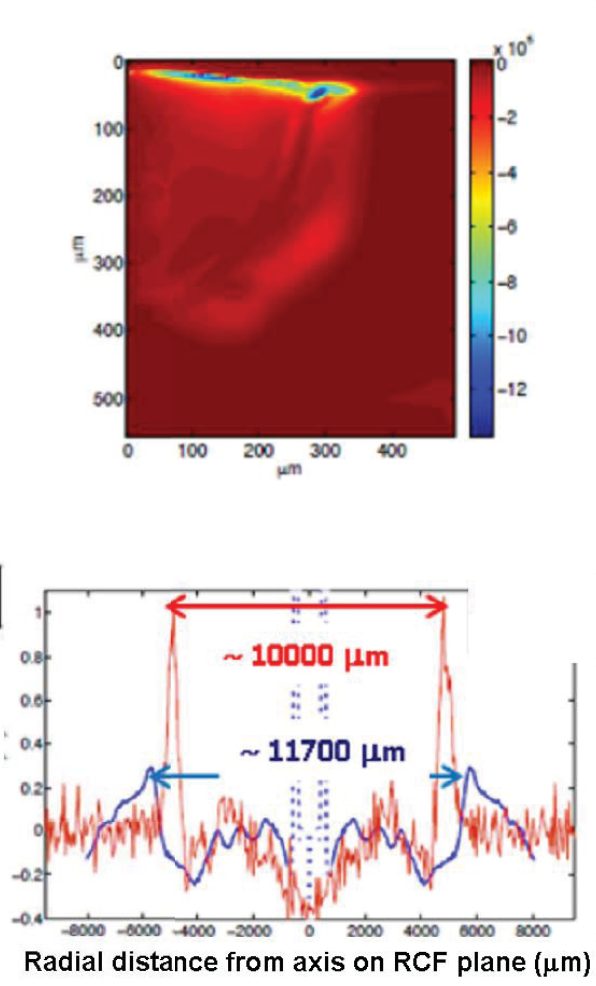
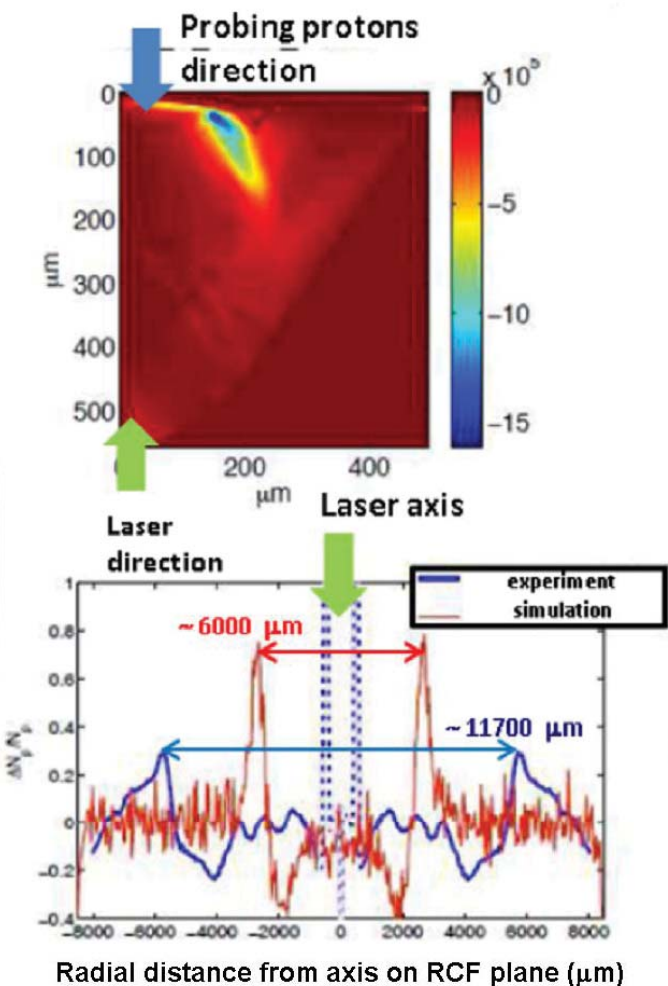
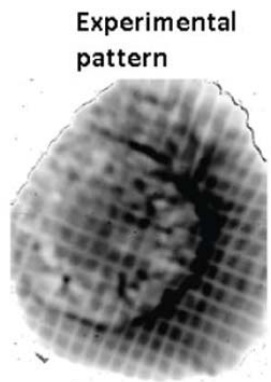
2D FP

Hydro-rad

# We could observe that NL heat transport plays an essential role to correctly model the B field

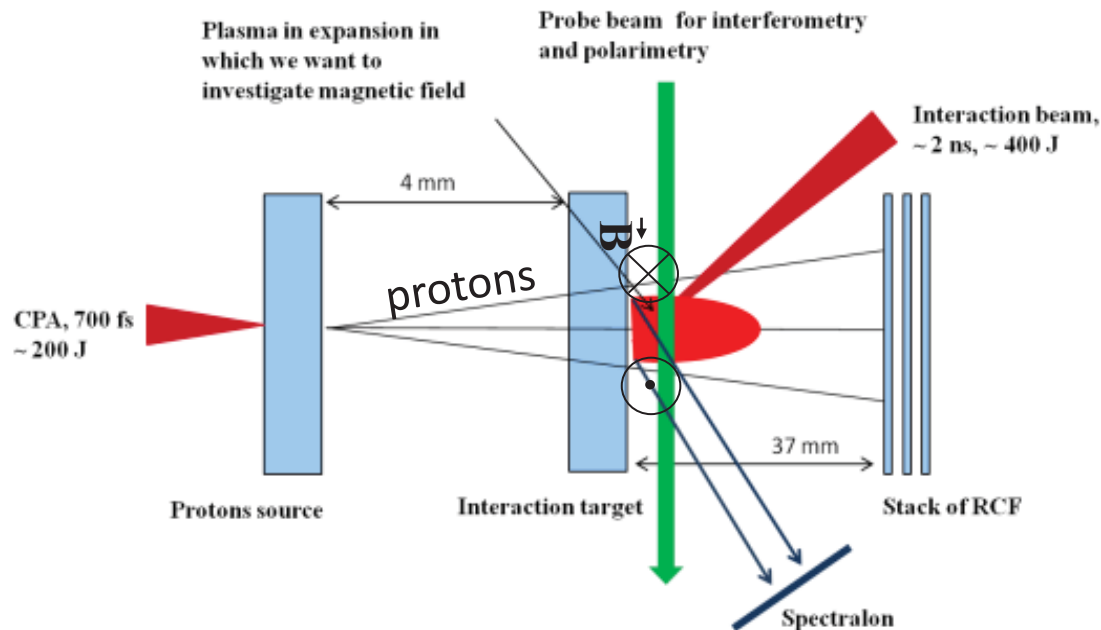
Local N+R-L

Nonlocal N+R-L



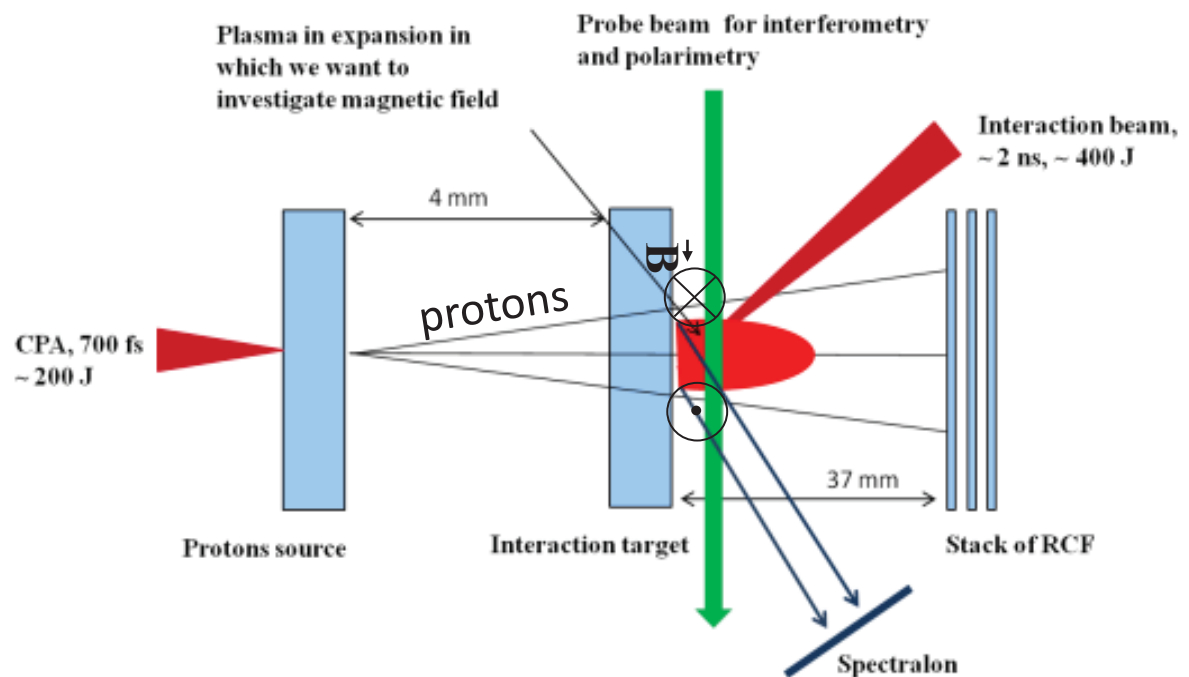
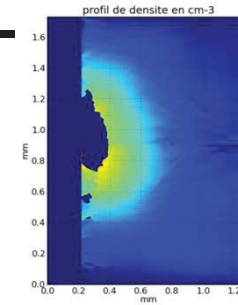
# But the late evolution of the B field is still pending

- The LULI2000 campaign probed the plasma only at the beginning of the high-power laser irradiation
- MHD could reproduce well the B-fields, but was validated against the kinetic model only over short times
- Li et al. (07, 09) showed issues with Masnex modeling of late time evolution of B fields
- This motivated us to study the long-term B evolution



# The experiment (@ Titan/LLNL) used many diagnostics to constrain the simulations

- Visible transverse interferometry ( $n_e$ )
- Visible transverse polarimetry ( $\mathbf{B}$  in the corona)
- Proton probing ( $\mathbf{B}$  in the dense plasmas)
- X-ray pinhole  
(time-integrated)
- Spectralon  
(scattered ns beam light)



## Hybrid simulations

Can hybrid code handle magnetic reconnection in HEDP ?

Physical hypotheses :

- ▶ Quasi-neutrality :  $n_e \sim n_i$  (but  $\nabla \cdot \mathbf{E} \neq 0$ )
- ▶ Electrons mobility  $\rightarrow \infty$  :  $\mathbf{V}_e$  is such as  
$$\mathbf{J} = ne(\mathbf{V}_i - \mathbf{V}_e) = \nabla \times \mathbf{B}$$
- ▶ Closure equation for electrons : isotherm or adiabatic
- ▶ Neglect transverse component of displacement current
- ▶ Hence, needs an Ohm's law :

$$\mathbf{E} = -\mathbf{V} \times \mathbf{B} + N^{-1}(\mathbf{J} \times \mathbf{B} - \nabla \cdot \mathbf{P}_e) + \eta \mathbf{J} - \eta' \Delta \mathbf{J}$$

$\Rightarrow$  Can solve  $kd_i \lesssim 1$ ,  $\omega/\Omega_i \gg 1$ , but no electron scales (neither spatial, nor temporal), and no plasma frequencies. Well suited if  $\omega_p/\Omega \gg 1$  (1000 in the solar wind).



## What is needed for reconnection ?

To trigger a reconnecting instability, one needs an electric field such as  $\nabla \times E_{\parallel} \neq 0$

⇒ If we are not interested in the onset (at electron scales), small (numerical) resistivity can do the job.

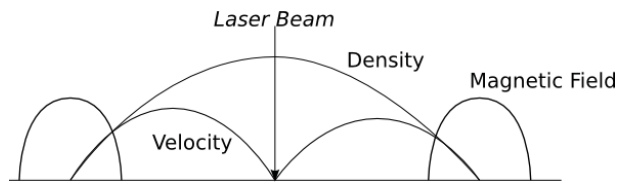
For collisionless &  $\beta \sim 1$  plasmas, GEM challenge (Birn 2001) showed that when the Hall effect is considered, the reconnection rate does not depend on the formalism.

⇒ What about HEDP ?

Nernst and Righi-Leduc effects are not considered. Collisions can be included.

## HECKLE is versatil, 3D & parallelized

Initial set up, very close to Fox et al. 2011 :



⇒ 2 bubbles initially (and not 2 halves) including a background :

- ▶ Can handle asymetries on  $\mathbf{B}$ ,  $n$ ,  $T$ ,  $\mathbf{V}$ ...
- ▶ Can handle non-coplanar configurations : set a given angle of rotation around the 2 directions of the target plane.

⇒ Plus few cautions to get  $\nabla \cdot \mathbf{B} = 0$  and periodic boundary conditions.



## Initial conditions

Old standing problem : if the initial set-up is not a kinetic equilibrium, energy has to be redistributed... generally in a wave.

This is the case if one uses a Maxwellian with an initial fluid equilibrium.

In our conditions, we are not even at an equilibrium... and of course we launch a  $k_{\perp}$  (magnetosonic) mode at  $t = 0$ .

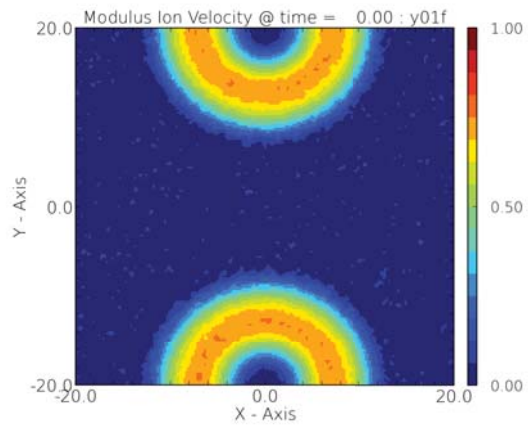
A kinetic equilibrium is generally very hard to find even in apparently simple topology : We do not think about it for laser configuration !

⇒ Is it a problem for the reconnection process ?



## Initial wave

This wave is very clear on the modulus of total ion velocity :

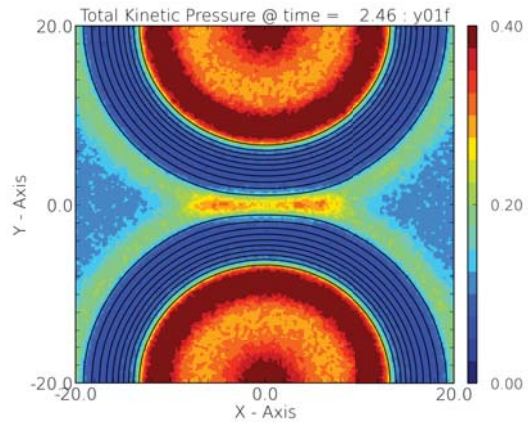


⇒ This question the initial set-up... that nevertheless seems coherent with hydro-radiative simulations. The associated RAM pressure is small.



## Initial Slow-down

The initial ion velocity is also slowed-down prior reconnection :

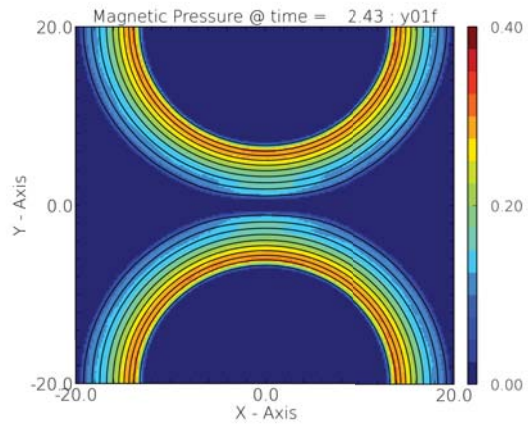


⇒ Essentially because of the enhancement of electron density at the noze (snowplow problem). Mostly a consequence of the background...



## Initial magnetic flux pile-up ?

Redistribution of magnetic pressure in the flux tube :

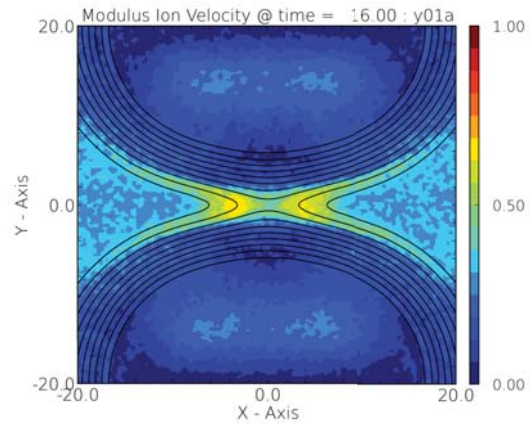


⇒ The flux tube is pushed from inner side, and blocked in the outer side. The total magnetic flux to be reconnected is conserved...



## First run

Let's have a look at reconnection...

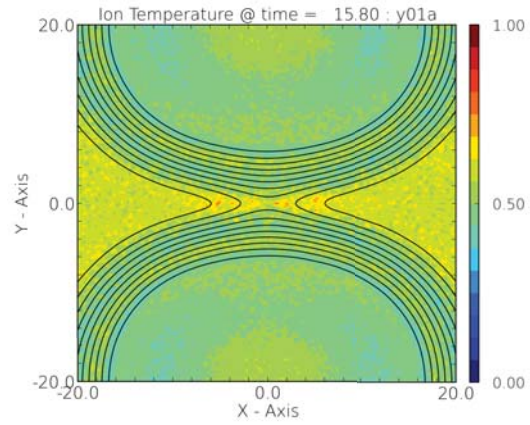


⇒ Fast reconnection with an outward jet velocity bounded by the Alfvén speed... as in GEM challenge.



## Energy budget

Is this accelerated flow associated to heating ?

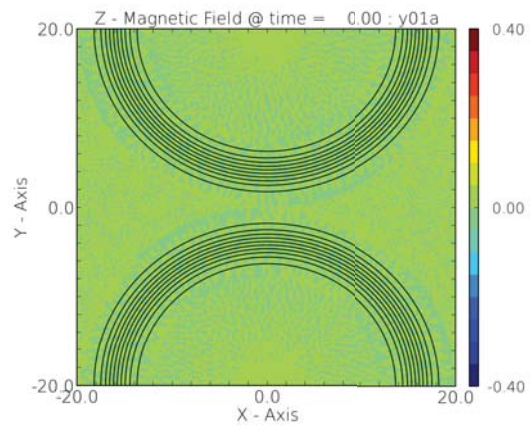


⇒ Energy budget is debated (Birn et al. 2010, Aunai et al. 2011), and should be measured in experiments.



## Fast reconnection ?

Are we talking about fast reconnection ?

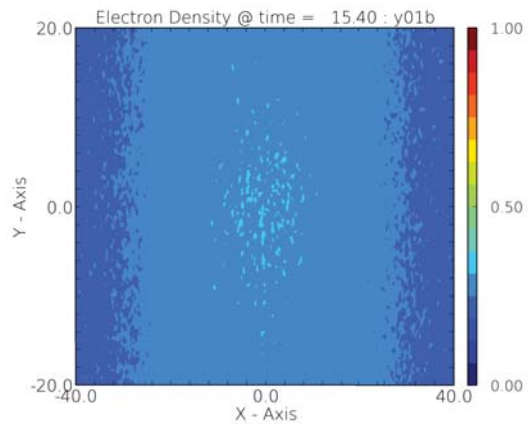


⇒ Yes. Electron jets at the separatrix, results in a Hall electric field : the Hall component of the magnetic field (Mandt et al. 1994)... should also be measured in experiments.



# Plasma dynamics with $B = 0$

If one doubts about magnetic reconnection in HEDP :



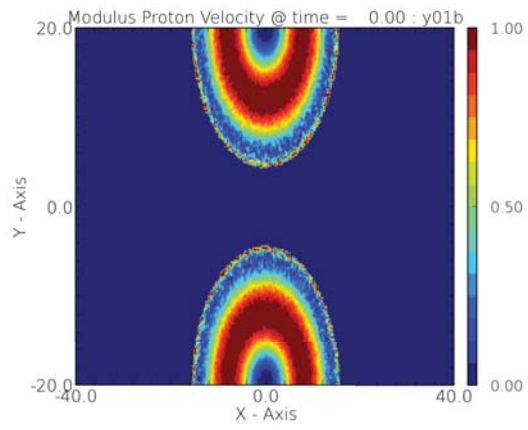
⇒ Totally diferent picture,





# Plasma dynamics with $B = 0$

If one doubts about magnetic reconnection in HEDP :

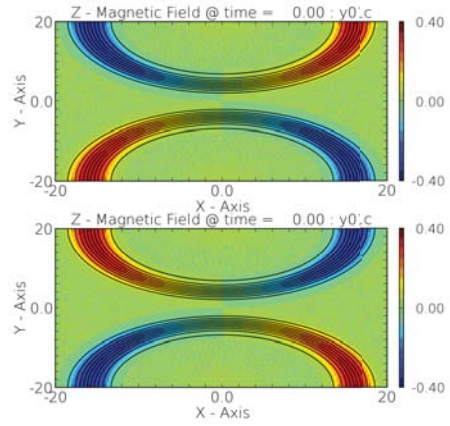


⇒ Totally diferent picture,



# Non-coplanar reconnection

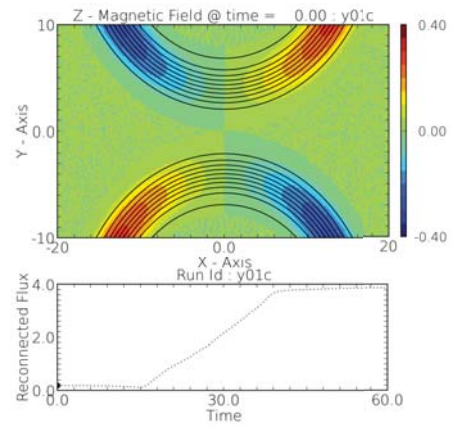
Depends on the angle : salient or reflex ?



⇒ Is the Hall magnetic field a cause or a consequence of magnetic reconnection ?



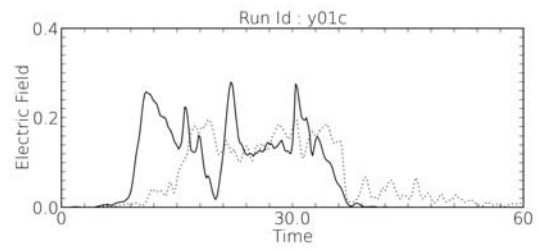
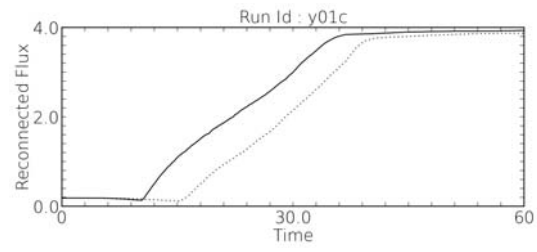
# Non-coplanar reconnection



⇒ No magnetic reconnection onset with the "wrong"  $B_Z$ .



## Consequence for the reconnection rate



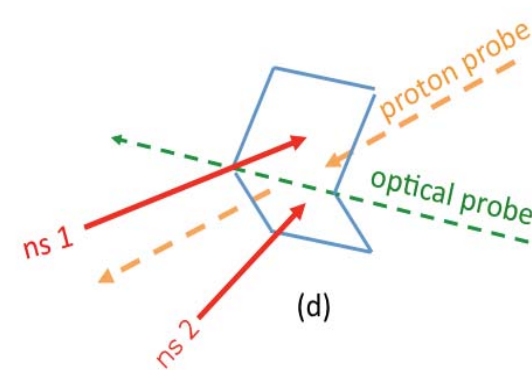
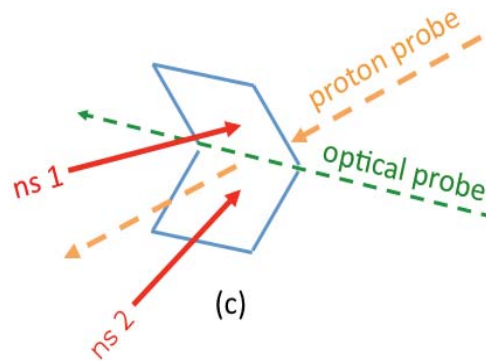
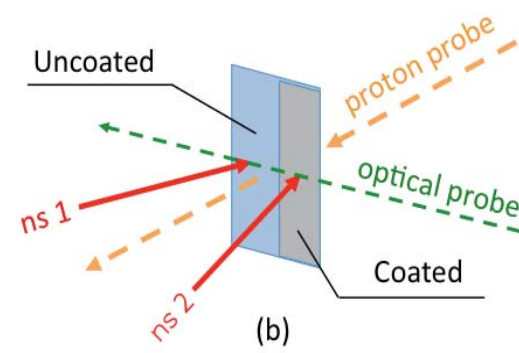
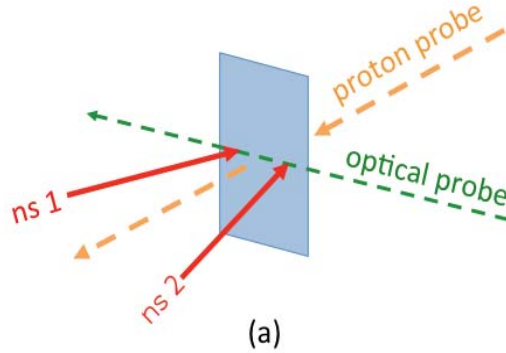
⇒ Reconnection is triggered later, but at a same rate.



# Short-term program on B-field reconnexion experiments

- A beamtime is planned at GSI in 2014 using Phelix

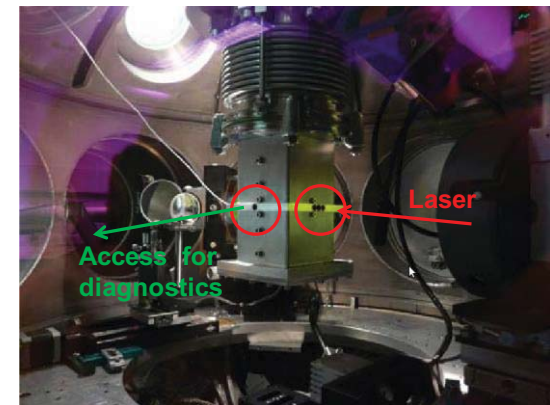
- Aim: to study symmetric (a) and non-symmetric (b)(c) and (d) reconnexion



# Long-term program

- Study macroscopic large-scale consequences of reconnection (solar corona)

- Aims:



- Importance of so-called guide field (use of pulsed-power unit we developed coupled to the laser)
- Identification of QSLs and study of energy release in slip-running reconnection (using the OHM code)